

## C-band Dual-polarimetric Radar Signatures of Hail

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### ABSTRACT

Many studies have demonstrated that dual-polarimetric radar is an effective tool for hail detection. Studies employing radar at S-band frequencies have typically found that hail is characterized by high reflectivity ( $Z_h > 50$  dBZ) and near zero differential reflectivity ( $-1 < Z_{dr} < 1$  dB) at and just above the ground. Several studies have hypothesized that hail signatures observed by higher frequency radars, such as C-band, should be similar. To test the hypothesis, the dual-polarimetric radar signatures of nine hail events, including 46 hail producing convective cells and 172 hail reports, are carefully documented using C-band radar observations over the Alabama and Tennessee border region. Unlike prior results at S-band, this study shows that hail at C-band is typically characterized by high  $Z_h$  ( $> 50$  dBZ) and high  $Z_{dr}$  (3 – 8 dB) at and just above the ground.

### 1. Introduction

Severe storms are important to identify and forecast because of the significant damage and fatalities that they can cause. A severe thunderstorm is defined by the National Weather Service (NWS) as a storm that produces large hail (diameter  $\geq 1.9$  cm at the start of this study), strong winds ( $> 26$  m/s), or a tornado. During the course of this study, the NWS definition of severe hail changed to a hail diameter  $\geq 2.5$  cm<sup>1</sup>. One large hail producing storm can cause 100's of millions of dollars of property damage to crops, livestock, homes, and automobiles (Hillaker et al. 1985, Changnon 1999). Large hail, in rare cases, can even cause bodily injury or death to people and livestock (Ludlam 1980, Appendix 8.12; National Climate Data Center

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<sup>1</sup> National Weather Service Instruction (NWSI) 10-511, April, 19, 2010: <http://www.nws.noaa.gov/directives/>

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(NCDC)). Clearly, it is important to identify areas of hail so that lives and property can be protected. To this end, ground-based radar is the most widely utilized observational tool for detecting and nowcasting hail (e.g. Greene and Clark 1972, Eccles and Atlas 1973, Mather et al. 1976, Bringi et al. 1984, Edwards and Thompson 1998).

Dual-polarization radar observations have improved the ability of the radar meteorologist to identify hydrometeor type, including hail. Studies for the classification of hydrometeors based on their polarimetric variables have been conducted at both S-Band (e.g. Aydin et al. 1986, Bringi et al. 1986, Hubbert et al. 1998, Straka et. al. 2000), C-Band (e.g. Meischner et al. 1991, Höller et al. 1994, Keenan et al. 2003, Deierling et al. 2005, Tabary et al. 2009a, Tabary et al. 2010) and X-band (e.g. Dolan and Rutledge 2009, Snyder et al. 2010) wavelengths. The use of C-band dual-polarimetric radars is becoming more common among worldwide operational weather services, including Australia and many Asian and European countries, the private market (e.g. media broadcasting) for forecasting in the United States, and worldwide universities for research purposes due to their affordability. As a result, it is important to document carefully the C-band dual-polarimetric radar signatures of hail. This study examines the polarimetric radar signatures of 46 well sampled, hail producing convective cells on nine different days over the Alabama and Tennessee border region using the UAHuntsville-NASA Advanced Radar for Meteorological and Operational Research (ARMOR) (Petersen et al. 2007).

## **2. Previous dual-polarimetric radar hail studies**

There are numerous literature examples demonstrating the effectiveness of hail identification using dual-polarimetric radar variables. For example, differential reflectivity ( $Z_{dr}$ )

is a measure of the decibel reflectivity difference between horizontal ( $Z_h$ ) and vertical ( $Z_v$ ) polarizations (Bringi and Chandrasekar 2001 p. 381). Therefore,  $Z_{dr}$  is helpful in estimating the oblateness of a hydrometeor. The more oblate the particle the larger  $Z_{dr}$  is. In rain,  $Z_{dr}$  is positive (typical values of 1 to 4 dB) and steadily increases with  $Z_h$  (Bringi and Chandrasekar 2001 p. 397). For large  $Z_h$  ( $> 50$  dBZ),  $Z_{dr}$  is typically  $> 2.5$  dB in rain. Hailstones are typically less oblate than raindrops and therefore are associated with lower  $Z_{dr}$ . The effect of oblateness on  $Z_{dr}$  in hail can be further reduced relative to rain because of the random orientations associated with tumbling and a lower dielectric constant for hail (Straka et al. 2000). As a result, hail can appear to the radar as an “effective sphere” and  $Z_{dr}$  is typically near 0 dB. Several S-band dual polarimetric radar-based studies have found hail to be characterized by near zero  $Z_{dr}$  ( $|Z_{dr}| < 1$  dB) and large  $Z_h$  ( $> 50$  dBZ) near the surface (Bringi et al. 1984, Illingworth et al. 1986, Bringi et al. 1986, Aydin et al. 1990, Herzegh and Jameson 1992, Hubbert et al. 1998). The dual-polarimetric radar-based hail signature of a minimum in  $Z_{dr}$  has been termed the “ $Z_{dr}$ -hole” (Wakimoto and Bringi 1988) because of the low  $Z_{dr}$  (hail) region surrounded by a higher  $Z_{dr}$  (rain) area in a region of large  $Z_h$  ([Fig. 1](#)) (Bringi et al. 1986, Bringi and Chandrasekar 2001, p. 451-456). Other radar studies at S-band have shown that  $Z_{dr}$  associated with melting hail can be moderately positive ( $0 < Z_{dr} < 2.5$  dB) but is typically smaller than in rain for a given  $Z_h$  (Leitao and Watson 1984, Aydin et al. 1986, Straka et al. 2000, Heinselman and Ryzhkov 2006, Depue et al. 2007). This anti-correlation between large  $Z_h$  and relatively small  $Z_{dr}$  forms the basis of most S-band dual-polarimetric radar hail detection algorithms. Several authors suggest that these methods to detect large hail can be directly applied at C-band (e.g. Vivekanandan et al. 1990, Aydin and Giridhar and 1992, Bringi and Chandrasekar 2001 p 451-452).

The correlation coefficient ( $\rho_{hv}$ ) is the co-polar correlation of the pulse to pulse returns between the horizontal and vertical polarizations in a given radar volume. In a radar volume that contains rain,  $\rho_{hv}$  is expected to be high, greater than 0.98 at S-band and 0.95 at C-band (e.g. Bringi et al. 1991, Carey et al. 2000, Keenan et al. 2000). The suppressed values of  $\rho_{hv}$  (0.95 to 0.98) in rain at C-band are associated with large drops that produce resonant scattering. Values of  $\rho_{hv}$  for hail at S-band (Balakrishnan and Zrnica 1990) and C-band (Tabary et al. 2010) have been found to be below 0.95. Low ( $< 0.95$ )  $\rho_{hv}$  in regions of hail and melting hail can be explained by non-Rayleigh (Mie or resonant) scattering, irregular shaped hydrometeors, mixture of hydrometeors (e.g. rain and melting hail), and a variety of hydrometeor shapes, dielectric constants and fall modes (Balakrishnan and Zrnica 1990).

At C-band frequency, Meischner et al. (1991), Ryzhkov et al. (2007), Tabary et al. (2009a) and Tabary et al. (2010) have shown that near 0 dB or even low ( $< 2.5$  dB)  $Z_{dr}$  is not necessarily what can be expected in areas of melting hail (Fig. 2). Meischner et al. (1991) attributes the anomalously high  $Z_{dr}$  ( $> 5$  dB) to smaller melting hail and resonance effects due to the smaller wavelength. Smaller melting hailstones can appear to the radar as large wet raindrops and exhibit high  $Z_{dr}$ . Resonance begins to take place at C-Band for rain and melting hailstones larger than about 5 mm (Zrnica et al. 2000). Therefore, when large particles are present  $Z_{dr}$  can be significantly higher at C-Band compared to S-Band (Zrnica et al. 2000). Ryzhkov et al. (2007) suggest that the anomalously high  $Z_{dr}$  can be explained by large raindrops and small melting hail that dominate  $Z_h$  in melting hail shafts. Tabary et al. (2009a) and Tabary et al. (2010) demonstrate that confirmed cases of melting hail and rain and hail mixtures are typically associated with  $Z_{dr} > 3$  dB and as high as 8 dB or more. All of the above C-band studies indicate that the S-band approach for hail identification of high  $Z_h$  and low  $Z_{dr}$  may not always work at C-

band. On the other hand, Bringi and Chandrasekar (2001) (p.451-452) and Ryzhkov et al. (2009) suggest that the  $Z_{dr}$ -hole ( $|Z_{dr}| < 1.0$  dB and large  $Z_h > 50$  dBZ) should be observed at C-band when hailstones are large and numerous enough to dominate  $Z_h$  and appear as effective spheres. Using a radar model of melting hailstones, Vivekanandan et al. (1990) suggest that hail above 12-15 mm in diameter shed their melt water and tumble, leading to the “ $Z_{dr}$ -hole” signature at C-band. Clearly, dual-polarimetric radar signatures of hail at C-band must be better understood in order to properly identify hail and provide warning of severe storms.

### **3. Data and Methodology**

The radar used for this study is the UAHuntsville-NASA ARMOR C-band dual-polarimetric radar, which is located at the Huntsville International Airport. Further specifications for ARMOR can be found in Petersen et al. (2007). For all but two events (May 20, 2008 and August 2, 2008), the ARMOR is in a PPI sector volume scanning mode, which includes multiple scanning elevation angles between user selected azimuths and is used to sample select storms well in the horizontal and vertical every 2 to 3 minutes. The other scanning method is a routinely used operational series of full  $360^\circ$  surveillance scans at the lowest three elevation angles ( $0.7^\circ$ ,  $1.3^\circ$ , and  $2.0^\circ$ ) that repeat every 3 to 5 minutes. ARMOR  $Z_h$  and  $Z_{dr}$  are the primary variables used to characterize hail in this study. ARMOR  $\rho_{hv}$  is used to identify potential areas of rain-hail mixtures associated with melting hail. At C-band, a mixture of partially melted hail and large rain drops originating from melting hail should be characterized by low  $\rho_{hv}$  ( $< 0.95$ ) because of resonant scattering and a variety of hydrometeor shapes, dielectric constants and fall modes in the radar resolution volume (Balakrishnan and Zrnicek 1990,

Meischner et al. 1991, Tabary et al. 2010).  $Z_{dr}$  has been calibrated using vertically pointing scans in light rain (Gorgucci et al. 1999).  $Z_h$  has been calibrated using self-consistency among dual-polarimetric variables (Ryzhkov et al. 2005).

$Z_h$  and  $Z_{dr}$  have been corrected for attenuation and differential attenuation, respectively, using the differential phase (Bringi et al. 2001). ARMOR is run for operational purposes across the Tennessee Valley and the data has been shown effective in other studies (e.g. Deierling et al. 2005, Gatlin et al. 2009). Recently, attempts have been made to develop and use improved propagation correction methods that account for “hot spots” associated with regions of resonant-sized large rain drops and/or melting hailstones that cause enhanced attenuation and differential attenuation at C-band (Carey et al. 2000, Tabary et al. 2009b, Gu et al. 2011). Although promising, these recent methods have not yet been widely tested for their usefulness with operational objectives and are therefore not utilized herein.

In order to investigate the C-band polarimetric signatures of hail, ARMOR data and storm reports are used from 9 different hail events, including 46 convective cells and 172 hail reports ([Table 1](#)). Hail report sizes range from 0.6 cm to 10.8 cm and average about 3.3 cm. The overwhelming majority of hailstone reports have sizes  $\geq 1.9$  cm, which met the NWS definition of severe during most of this study, while 65% of the reports (111) meet the current NWS definition of severe hail. The locations, magnitudes, and timing of observed hail were obtained from the NCDC *Storm Data*. For the prolific hail producing event of April 10, 2009, an Internet survey of hail properties was also conducted among UAHuntsville and NASA staff and students. Each hail report was evaluated carefully for accuracy by comparing archived radar data to its time and location. If a 50 dBZ echo did not occur within +/- 15 minutes and 1 km of a hail report, then the report was not used in the study. The 1 km/15 minute criteria were used to

capture all potential hail data and to account for inaccuracies in the location and time of reports. [Table 2](#) shows different environmental factors that affect the hailstone melting process such as the melting level, temperature lapse rate, relative humidity and the surface temperature. The data were obtained from archived RUC (Rapid Update Cycle) model soundings from the Huntsville, AL area.

Although plan position indicator (PPI) images of  $Z_h$  and  $Z_{dr}$  for all 46 hail producing convective cells were visually inspected for hail signatures (e.g., [Fig. 3](#)), the quantitative results in this study rely on characterizing  $Z_h$  and  $Z_{dr}$  within a 1 km radius of each quality controlled hail report. Analysis radii of 0.5 km and 1.5 km were also tested but had little effect on the overall results and are therefore not shown. Joint absolute frequency histogram plots of  $Z_{dr}$  versus  $Z_h$  at low levels (i.e., data below 1 km in height) and contoured frequency by altitude diagrams (CFADs) of  $Z_{dr}$  versus height (Yuter et al. 1995) are used to evaluate hail signatures at C-band for the April 10, 2009 event (5 cells, 77 hail reports). For the composite analysis of all nine hail events (46 cells, 172 hail reports), joint relative frequency histograms of  $Z_{dr}$  versus  $Z_h$  are created using low-level data (data below 1 km in height) in which the absolute count in each bin is divided by the total bin count and is expressed in percent. Similarly, CFADs of the  $Z_{dr}$  and  $\rho_{hv}$  versus height relative to the 0 °C level ( $H_0$ , km) were used to evaluate the signature of melting hail associated with the composite analysis. These diagrams are also made by dividing the absolute count in each bin by the total bin count and are expressed in percent. Histogram bin sizes used in this study are: 2 dBZ for  $Z_h$ , 0.5 dB for  $Z_{dr}$ , 0.01 for  $\rho_{hv}$ , and 250 m for height.

## 4. Results

We begin by presenting results from the prodigious hail producing event of April 10, 2009. Although not shown herein, the polarimetric signatures of the other hail events have been studied separately and are consistent with the results from April 10, 2009. To summarize the overall results of the study, we show composite statistical analyses of all nine hail events.

### *a. April 10, 2009 hail event*

The PPI shows a storm in close proximity to the radar (10-15 km) crossing over highly populated areas where a large number of reports were received ([Fig. 3](#)). The first PPI ([Fig. 3a](#)) is taken from the lowest elevation scan of  $0.7^\circ$  and is a good representation of  $Z_h$  and  $Z_{dr}$  closest to the surface. The  $Z_h$  shows areas in the cell of high values ( $> 55$  dBZ) associated with particles of large diameters possibly large raindrops, hail, and melting hail that fall into the resonance size region. It can be seen that high  $Z_h$  ( $> 50$  dBZ) is associated with high  $Z_{dr}$  (3-8 dB) near the surface ([Fig. 3a](#)). In some radar gates, the corrected  $Z_{dr}$  is as high as 10 dB. There is no evidence of the “ $Z_{dr}$ -hole” that is typically found at S-band. The PPIs in [Fig. 3b](#) indicate a persistent, large area of enhanced  $Z_{dr}$  (3-8 dB) associated with high  $Z_h$  ( $> 50$  dBZ) at heights of about 1.1 km above radar level (i.e., range  $\sim 15$  km and elevation angle =  $4.2^\circ$ ). Since the  $0^\circ\text{C}$  height level is at 3 km, it is hypothesized that melting of resonant sized hailstones and a mixture of hydrometeors is causing this high  $Z_{dr}$  (3-8 dB). The PPIs in [Fig. 3c](#) at  $11^\circ$  elevation angle show that higher in the storm (i.e., height  $> 3$  km at range  $> 15$  km) low  $Z_{dr}$  (-1 to 1 dB) is associated with high  $Z_h$  ( $> 50$  dBZ). The low  $Z_{dr}$  is likely due to tumbling dry hail and graupel found above the  $0^\circ\text{C}$  height level.



The joint frequency histogram of  $Z_{dr}$  versus  $Z_h$  (Fig. 4) from the radar range gates located at heights below 1 km and within 1 km of the hail reports are characterized by anomalously high  $Z_{dr}$  for areas of high  $Z_h$ . There is a  $Z_{dr}$  mode of 5 dB that occurs at 54 dBZ. Maximum  $Z_{dr}$  values are observed as high as 10 dB (Fig. 4). For the  $Z_h$  bin from 55-57 dBZ, the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles of  $Z_{dr}$  are 4.0, 5.6, and 7.0 dB, respectively, with a mean of 5.4 dB. This indicates that 50 % of the  $Z_{dr}$  data within the 55 – 57 dBZ bin can be found between 4.0 and 7.0 dB. These results are similar to what was found in Meischner et al. (1991). However, the anomalously high  $Z_{dr}$  observed in Meischner et al. (1991) is hypothesized to be due to smaller melting hailstones that appear as large raindrops to the radar. When melting occurs, the water torus on the outside of the drop can become sufficient enough to stabilize the hailstone (Rasmussen and Heymsfield 1987). In this case, one would expect high  $Z_{dr}$ . Zrnic et al. (2000) notes that wet particles over 5 mm at C-band are in the resonant scattering regime and can cause peaks in  $Z_{dr}$  for particles greater than 5 mm. All reported hailstone diameters in this study are well over 5 mm, therefore, resonant scattering could be contributing to the anomalously high  $Z_{dr}$ . However, there is no certainty that only hail and melting hail are present in the data set and other hydrometeors, such as large raindrops, could have contributed to the signatures.

In addition to analyzing the  $Z_{dr}$  and  $Z_h$  polarimetric signatures, the hail signal  $H_{dr}$  developed by Aydin et al. (1986) for use at S-band was plotted (Fig. 4). Aydin and Giridhar (1992) suggest that this method can be applied to C-band with little modification. The solid red line in Fig. 4 is where  $H_{dr}$  is 0 dB. Areas to the right and below the line are where  $H_{dr}$  is positive and is taken to signify hail, while areas to the left and above the line are where  $H_{dr}$  is negative and is assumed to indicate rain (Aydin et al. 1986). It can be seen that just the opposite occurs in the C-band ARMOR observations. Most of the hail observations fall on the negative side of the

plot where rain is expected at S-band except for the points greater than 60 dBZ where Aydin et al. (1986) assumes data above this value are associated with hail. These results indicate that the conventional behavior of S-band  $Z_h$  and  $Z_{dr}$  does not apply at C-band for this hail event.

A striking result that also goes against traditional S-band knowledge can be found in the CFAD of  $Z_{dr}$  (Fig. 5). The 0 °C level indicated on the 17 UTC sounding from Redstone Arsenal, located 8 km to the east of the ARMOR C-band polarimetric radar, is around 3 km.  $Z_{dr}$  in this example is around 0 dB above the 0 °C level, which is what has also been observed at S-Band. As the hail continues below the 0 °C level,  $Z_{dr}$  begins to increase toward high positive values with a mode near the surface of 5.5 dB and a maximum of near 10 dB, which is unlike past S-band studies. For the height bin from 0 to 250 m, the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles of  $Z_{dr}$  are 4.0, 5.5, 6.9 dB, respectively, with mean of 5.5 dB while well above the 0 °C level at the 4.0 to 4.25 km bin the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles of  $Z_{dr}$  are -0.2, 0.1, and 0.5, respectively, with a mean of 0.2 dB. This indicates that near the surface (0 – 250 m) 50 % of the  $Z_{dr}$  data occurs between 4.0 and 6.9 dB while aloft (4.0 – 4.25 km) the distribution of data is narrower with the 50 % of data occurring between -0.2 and 0.5 dB. Therefore, near the surface resonant sized (> 5 mm at C-band), stably oriented and melting hailstones, possibly mixed with large raindrops, are likely enhancing  $Z_{dr}$ . The increase in  $Z_{dr}$  takes place below about 1.6 km where the hailstones have likely had enough time to begin melting and developing a sufficient water torus. Interestingly, the  $Z_{dr}$  distribution starts broadening, including more negative and positive values, at about 2 km before shifting to more distinctly positive values between 1.6 km and 1 km. Table 2 indicates mid-level average relative humidity for the April 10<sup>th</sup> event to be near 14 %. Rasmussen and Pruppacher (1982) indicate that atmospheres with lower RH's can delay the

onset of melting. This effect of low RH could be the reason why shifts in the  $Z_{dr}$  data for [Fig. 5](#) are seen near 1 km below the 0 °C level.

*b. Composite of all nine hail events*

The joint relative frequency distribution of  $Z_{dr}$  versus  $Z_h$  reveals a high  $Z_{dr}$  mode of 5 dB with  $Z_h$  between 51 to 57 dBZ ([Fig. 6](#)). [Figure 6](#) indicates the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles for the  $Z_h$  bin from 55 to 57 dBZ are 4.0, 5.5, and 6.9 dB, respectively, with a mean of 5.3 dB. Therefore, 50 % of the data in the 55 to 57 dBZ bin is in between 4.0 to 6.9 dB. In addition to the high  $Z_{dr}$  mode, an increase in  $Z_{dr}$  occurs as  $Z_h$  increases. The high  $Z_{dr}$  may be due to melting, stably oriented hailstones, possibly mixed with raindrops, which fall into the resonance regime at C-band (Meischner et al. 1991, Vivekanandan et al. 1990, Ryzhkov et al. 2007).

The CFAD of  $Z_{dr}$  indicates a clear shift in  $Z_{dr}$  from near 0 dB aloft (i.e., heights above 0 °C level) to large positive  $Z_{dr}$  values below the 0 °C level ( $H_0$ ) with the dramatic increase in  $Z_{dr}$  beginning at  $H_0 = -1.5$  km ([Fig. 7](#)). [Figure 7](#) is similar to [Fig. 5](#) due to the overwhelming majority of points from the April 10, 2009 event. For the  $H_0$  bin from -2.50 to -2.25 km, the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles for  $Z_{dr}$  are 3.2, 5.2, and 6.9 dB, respectively, with a mean of 5.3 dB while above the 0 °C level from  $H_0 = 0.75$  to 1.00 km bin the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles for  $Z_{dr}$  are -0.2, 0.1, and 0.5 dB, respectively, with a mean of 0.4 dB. [Figure 7](#) suggests that significant microphysical changes, which are likely driven by the melting of hail, are occurring to cause the sudden increase in  $Z_{dr}$  near the surface. [Figure 8](#) confirms this mixture below the freezing level by showing a dramatic shift in the  $\rho_{hv}$  data as well. For the  $H_0$  bin from -2.50 to -2.25 km, the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles for  $\rho_{hv}$  are 0.88, 0.93, and 0.96, respectively, with a mean of 0.92 while above the 0 °C level from  $H_0 = 0.75$  to 1.00 km bin the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup>

percentiles for  $\rho_{hv}$  are 0.95, 0.98, and 0.99, respectively, with a mean of 0.96. [Table 2](#) indicates a range of environmental factors that can contribute to melting hailstones. Mid-level relative humidity (RH) varies from 14 – 86 %. Rasmussen and Heymsfield (1987b) indicate that an atmosphere with higher RH's will experience melting first while hailstones falling in atmospheres with lower RH's will experience delayed melting closer to the surface. With the numerous hail reports from the April 10, 2009 event, this delayed melting associated with lower RH may be why the shift in the polarimetric signatures associated with melting are not observed until 1 km below the 0 °C height level (i.e.,  $H_0 = -1$  km). As seen in [Table 2](#), the heights of the 0 °C level range from 2.64 to 4.81 km above ground level (AGL). These 0 °C heights should provide significant time for melting to occur with surface temperatures ranging from 14.1 - 30.9 °C. Lapse rates for the events studies range from 5.1 to 7.6 °C/km. Rasmussen and Heymsfield (1987b) noted that hailstones can melt more quickly in higher temperature lapse rates.

Overall, the April 10, 2009 and composite analysis of  $Z_{dr}$  versus  $Z_h$  and  $Z_{dr}$  versus height reveal similar results. In regions of hail and melting hail, near surface values of  $Z_{dr}$  are typically 3 to 8 dB while  $Z_{dr}$  values above the 0 °C height level,  $H_0$ , are typically -0.2 to 0.5 dB. The transition from near zero  $Z_{dr}$ , which is consistent with dry, tumbling hail, to large  $Z_{dr}$ , which is consistent with resonance sized melting hail likely mixed with large rain drops, begins to occur around 1 km below  $H_0$  ( $H_0 = -1$  km). Composite analysis of  $\rho_{hv}$  with  $H_0$  reveals suppressed values of  $\rho_{hv}$  ( $< 0.95$ ) increasing in frequency between  $H_0 = 0$  km and  $H_0 = -1$  km and being quite common for  $H_0 \leq -1$  km, supporting the idea that the melting of hail is actively occurring between  $H_0 = 0$  km and  $H_0 = -1$  km and that a mixture of resonance sized melting hail and large rain drops is likely present at  $H_0 < -1$  km.

## 5. Summary and Conclusions

ARMOR observations show that the C-band dual-polarimetric radar-based signatures of large hail (average size of 3.3 cm) are typically characterized by high  $Z_h$  ( $> 50$  dBZ), high  $Z_{dr}$  (3-8 dB), and low  $\rho_{hv}$  ( $< 0.95$ ) at low-levels. Therefore, the hypothesis that C-band radar signatures of large hail are usually consistent with the S-band “ $Z_{dr}$ -hole” signature of high  $Z_h$  ( $> 50$  dBZ) and low  $Z_{dr}$  (-1 to 1 dB) at low-levels has been rejected through careful observation of nine hail events, 46 convective cells and 172 hail reports, suggesting that S-band techniques for identifying large hail using  $Z_{dr}$  cannot be directly applied at C-band without modifications.

It is hypothesized in this study along with others (Vivekanandan et al. 1990, Meischner et al. 1991, Zrníc et al. 2000, Ryzhkov et al. 2007, Tabary et al. 2009a, Tabary et al. 2010) that the high  $Z_{dr}$  (3-8 dB) signature exists due to melting, resonant-sized, stably and horizontally oriented (i.e., with major axis in the horizontal) hailstones and large raindrops that are present at low-levels. The vertical profile of  $\rho_{hv}$  confirms this hypothesis by showing clear lowering below the height of the 0 °C level. Additionally, the environmental parameters presented in [Table 2](#) show that melting hail is most likely present for each of these events near the surface. The clear lowering in  $\rho_{hv}$  ( $< 0.95$ ) associated with resonance has been shown in the studies above to cause enhanced  $Z_{dr}$  (3 - 8 dB) for wet particles over 5 mm. In addition, melting may play a role in setting up the right conditions for resonance including the necessary size, shape, fall mode, and dielectric constant.

By comparing collocated S-band (insignificantly attenuated) to C-band (potentially highly attenuated) frequency radar observations, Borowska et al. (2011) demonstrate that in the case of strong hail bearing storms, phase-based propagation correction methods that estimate

mean corrections coefficients over the entire storm path, such as the Bringi et al. (2001) attenuation correction algorithm used herein, can under estimate the attenuation and differential attenuation in and beyond “hot spots” associated with melting hail and large rain drops. If this is the case, then even higher  $Z_h$  and  $Z_{dr}$  would be expected in and down range of hail producing regions or hot spots. In other words, the fundamental conclusion of this study would likely not change even if the enhanced propagation effects of these hot spots were individually addressed – i.e., melting hail near the surface at C-band is typically characterized by high  $Z_h$  ( $> 50$  dBZ), high  $Z_{dr}$  (3 – 8 dB or more), and low  $\rho_{hv}$  ( $< 0.95$ ). Nonetheless, more research is required to test the proposed operational propagation correction methods of Tabary et al. (2009b) and Gu et al. (2011), which deal specifically with hot spots at C-band, and verify the results of this study. Additionally, more C-band radar observations in a variety of meteorological environments along with thorough in-situ hydrometeor documentation, as begun in Tabary et al. (2009a) and (2010), are required to refine and confirm these results.

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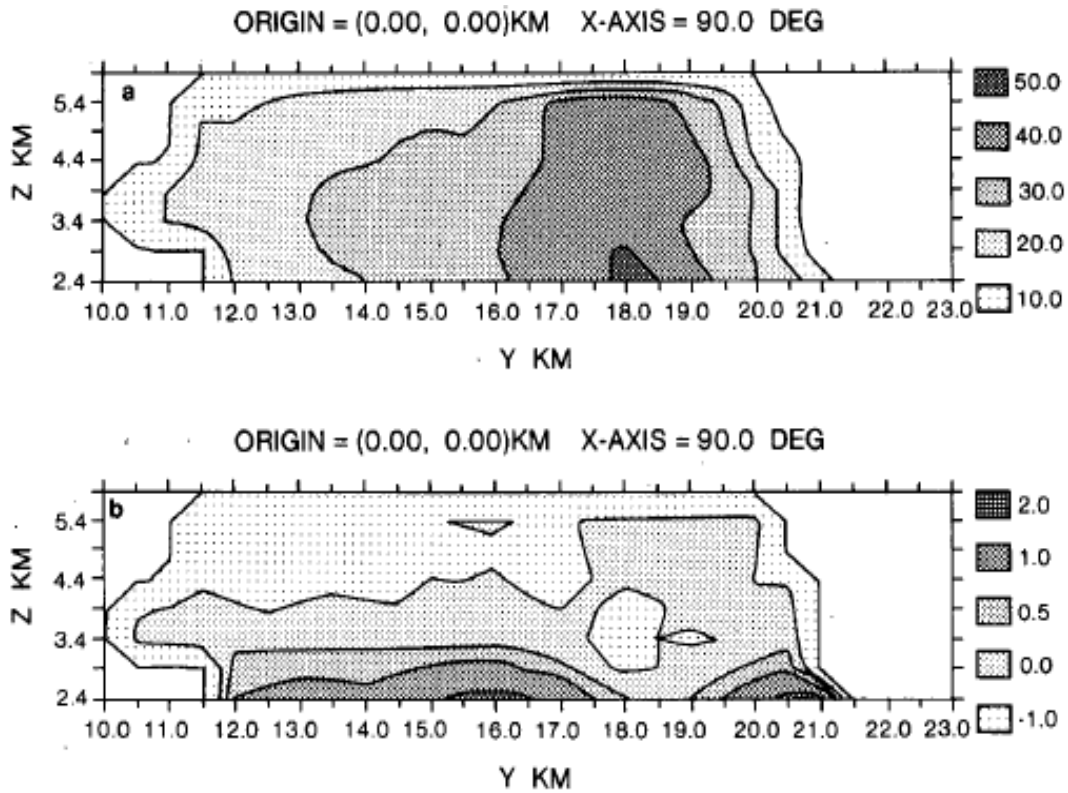
## TABLES AND FIGURES

**Table 1.** Overview of Hail Events

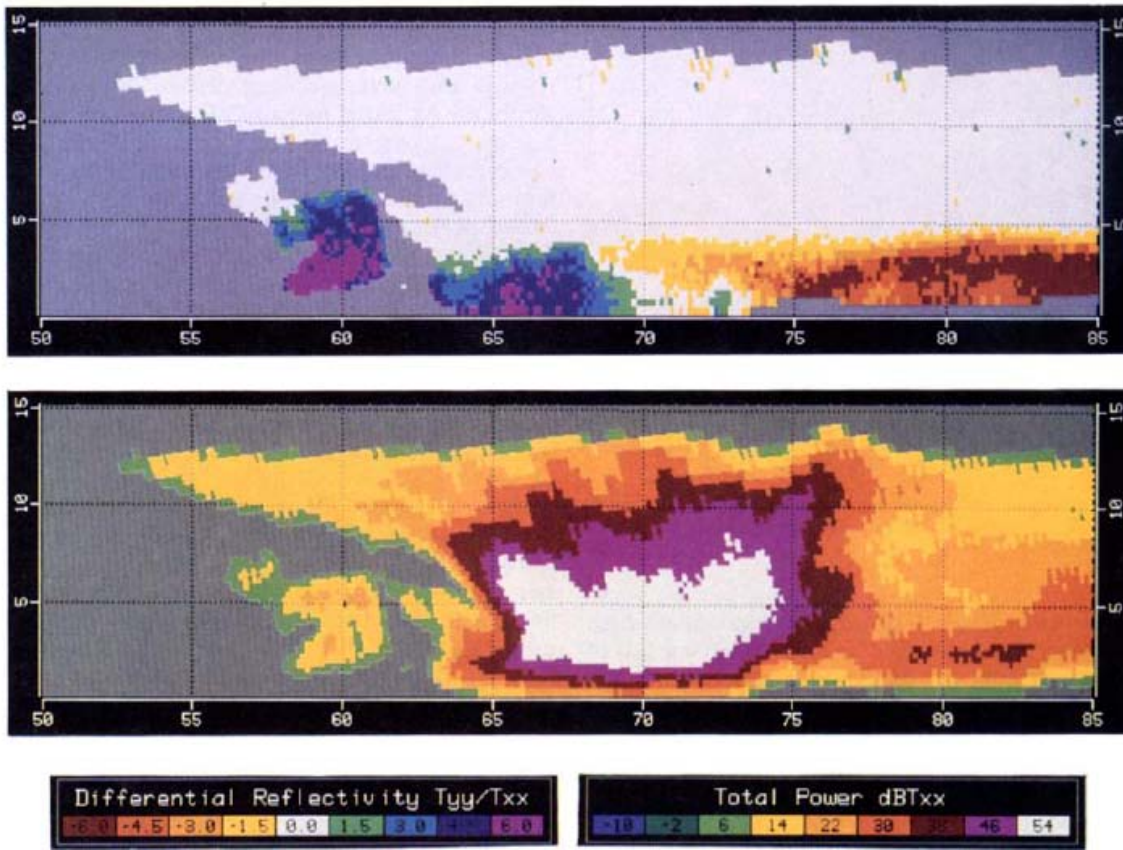
<i>Date</i>	<i>Number of Cells</i>	<i>Number of Hail Reports</i>	<i>Average Hail Size</i>	<i>Hail Size Range (Minimum to Maximum)</i>
February 21, 2005	9	25	2.4 cm	1.9 – 4.5 cm
April 3, 2007	2	9	2.4 cm	1.9 – 4.5 cm
March 15, 2008	4	7	2.1 cm	1.9 – 2.2 cm
August 2, 2008	4	19	3.3 cm	1.9 – 4.5 cm
March 28, 2009	4	4	2.5 cm	1.9 – 3.2 cm
April 10, 2009	5	77	4.0 cm	0.6 – 10.8 cm
April 13, 2009	3	4	2.3 cm	1.9 – 2.5 cm
January 21, 2010	2	6	3.3 cm	1.3 – 7.0 cm
March 12, 2010	10	21	2.9 cm	1.9 – 4.5 cm
<b>Total:</b>	46	172	3.3 cm	0.6 – 10.8 cm

**Table 2.** Overview of Environmental Parameters

<i>Date</i>	<i>Height of 0 °C</i>	<i>Lapse Rate (700-500 mb)</i>	<i>Average Mid-Level RH (700-500 mb)</i>	<i>Surface Temperature</i>
February 21, 2005	3.17 km	6.0 °C/km	43 %	21.1 °C
April 3, 2007	3.70 km	6.8 °C/km	33 %	26.4 °C
March 15, 2008	3.21 km	7.4 °C/km	30%	17.2 °C
August 2, 2008	4.81 km	6.3 °C/km	49 %	30.9 °C
March 28, 2009	3.22 km	5.5 °C/km	65 %	21.2 °C
April 10, 2009	2.99 km	7.6 °C/km	14 %	21.2 °C
April 13, 2009	2.81 km	5.1 °C/km	74 %	20.9 °C
January 21, 2010	2.64 km	6.8 °C/km	76 %	18.5 °C
March 12, 2010	2.85 km	7.2 °C/km	86 %	14.1 °C
<b>Average:</b>	3.27 km	6.5 °C/km	52 %	21.3 °C

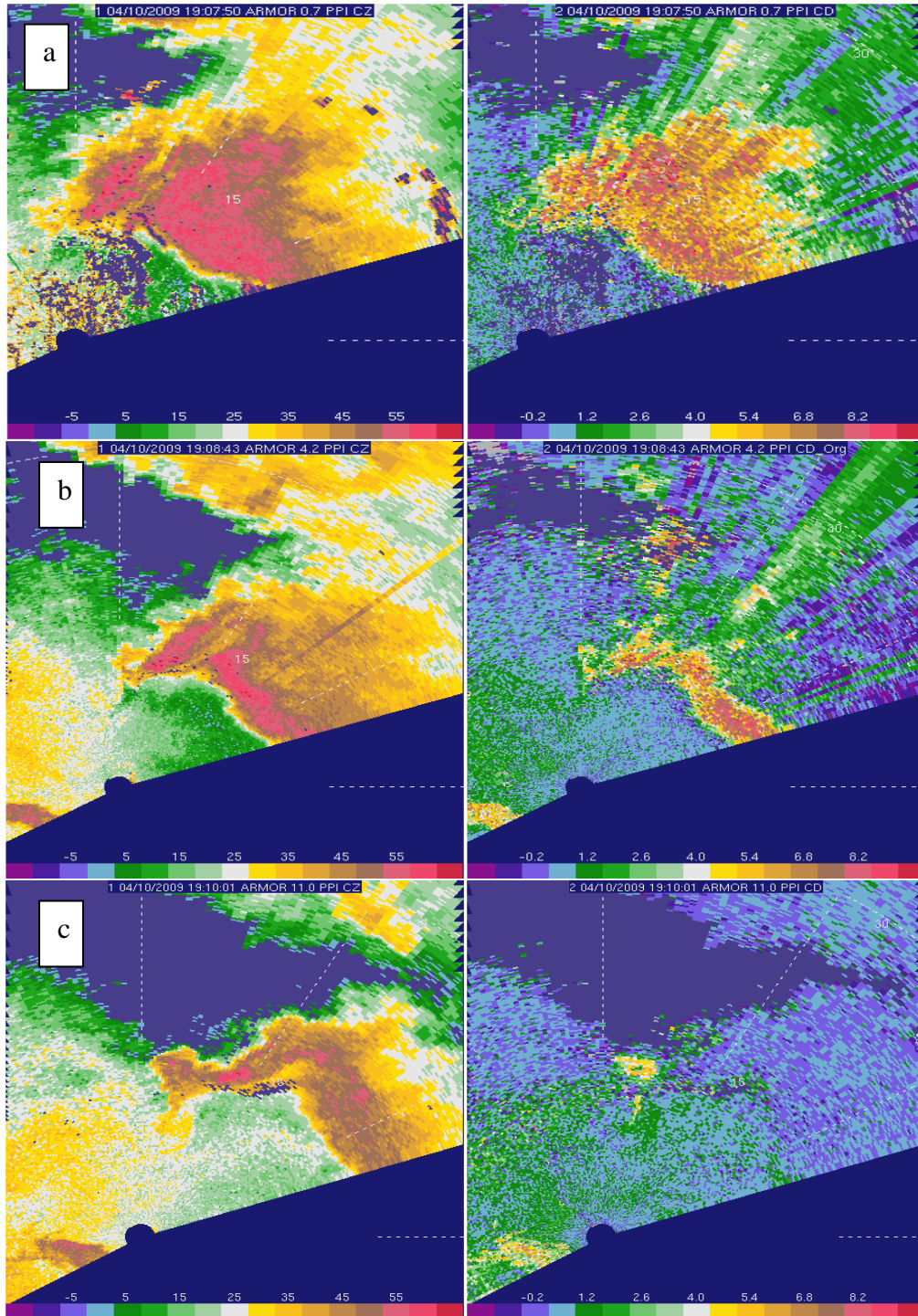


**Figure 1.** Adapted from Bringi et al. (1986). (a) Range Height Indicator (RHI) of reflectivity and (b)  $Z_{dr}$  in contours through a hail producing convective cell at S-band. The dual polarimetric radar-based hail signature, or “ $Z_{dr}$ -hole”, is observed between 18 and 19 km. (c) American Meteorological Society. Reprinted with permission.

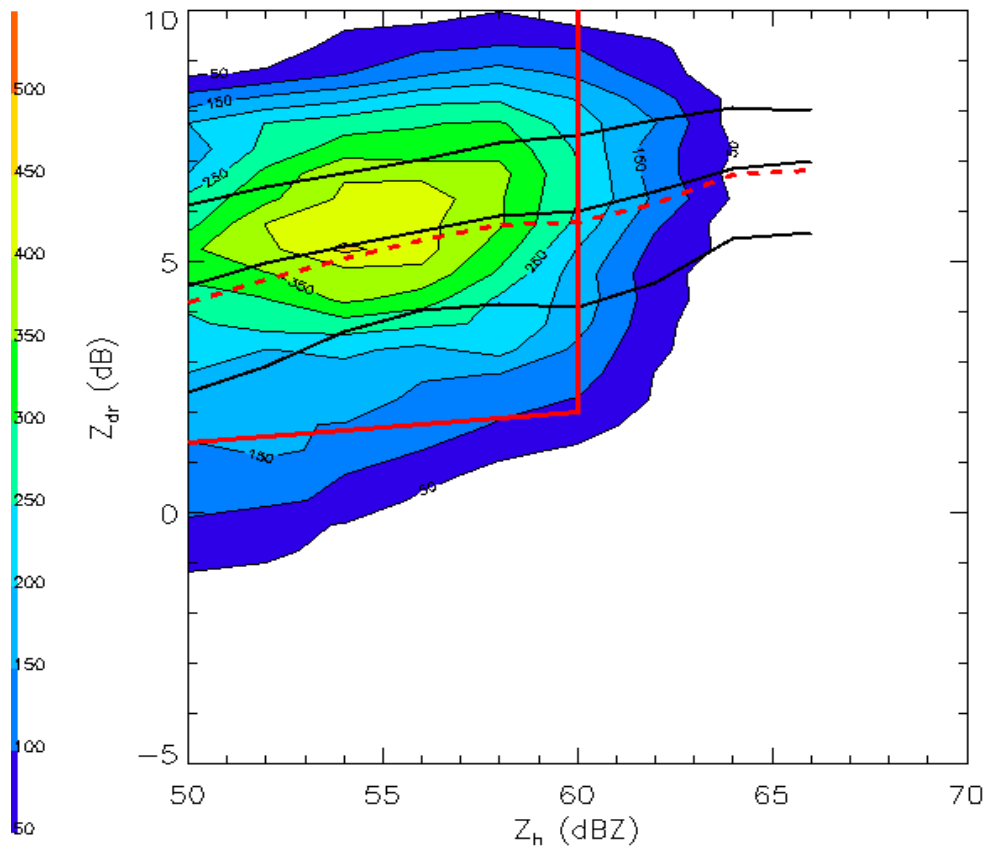


**Figure 2.** Adapted from Meischner et al. (1991). Vertical cross section (RHI) of  $Z_{dr}$  (top panel) and  $Z_h$  (bottom panel) through a hailstorm at C-band. High  $Z_{dr}$  ( $> 5$  dB) associated with high  $Z_h$  ( $> 45$  dBZ) can be seen between 65 and 70 km. (c) American Meteorological Society. Reprinted with permission.

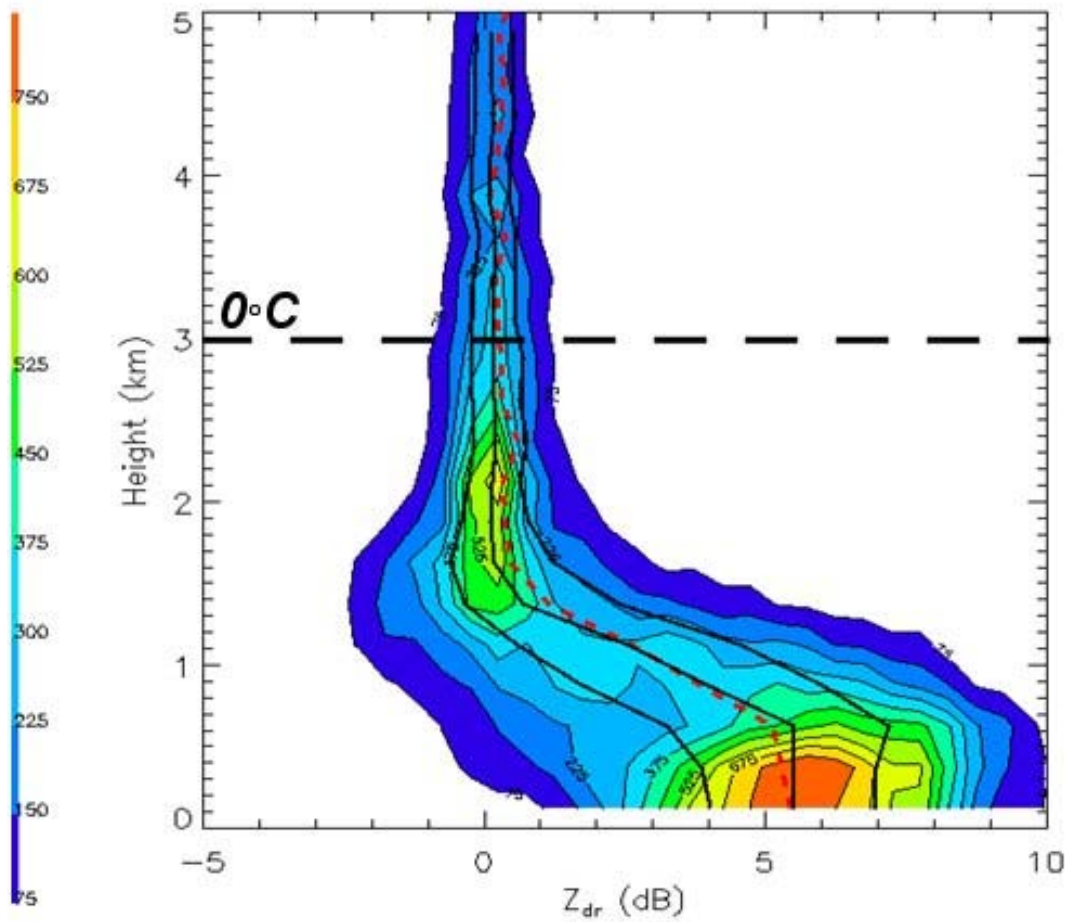




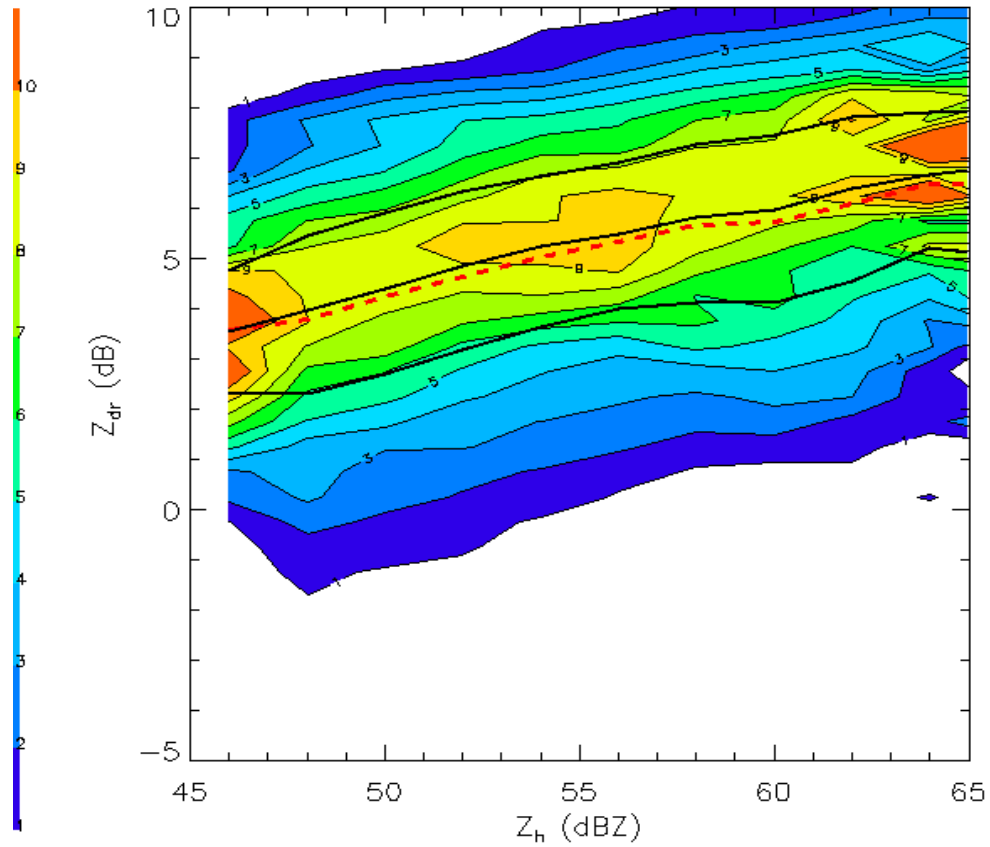
**Figure 3.** ARMOR Plan Position Indicator (PPI) view of  $Z_h$  (dBZ) (left) and  $Z_{dr}$  (dB) (right) on April 10, 2009 at 1907 UTC at (a)  $0.7^\circ$ , (b)  $4.2^\circ$ , and (c)  $11.0^\circ$  elevation angles. Horizontal east-west distance from ARMOR (km) versus north-south distance from ARMOR is shown.



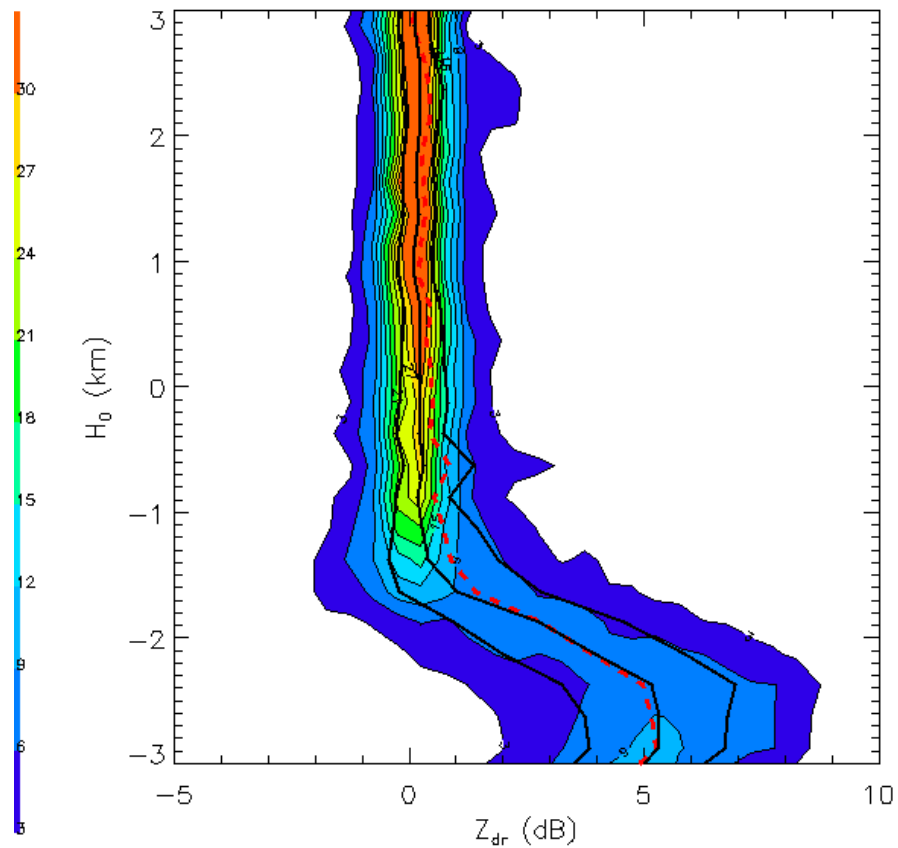
**Figure 4.** Joint frequency histogram of ARMOR  $Z_{dr}$  versus  $Z_h$  at low levels for April 10, 2009. All ARMOR data are taken from the lowest 1 km in height above the surface and from within 1 km horizontal distance of a hail report for values of  $Z_h > 45$  dBZ. The  $Z_h$  data are in 2 dB bins and the  $Z_{dr}$  data are in 0.5 dB bins. The thick black lines from bottom to top are the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles. The dashed red line is the mean. The solid red line is where  $H_{dr} = 0$  dB.



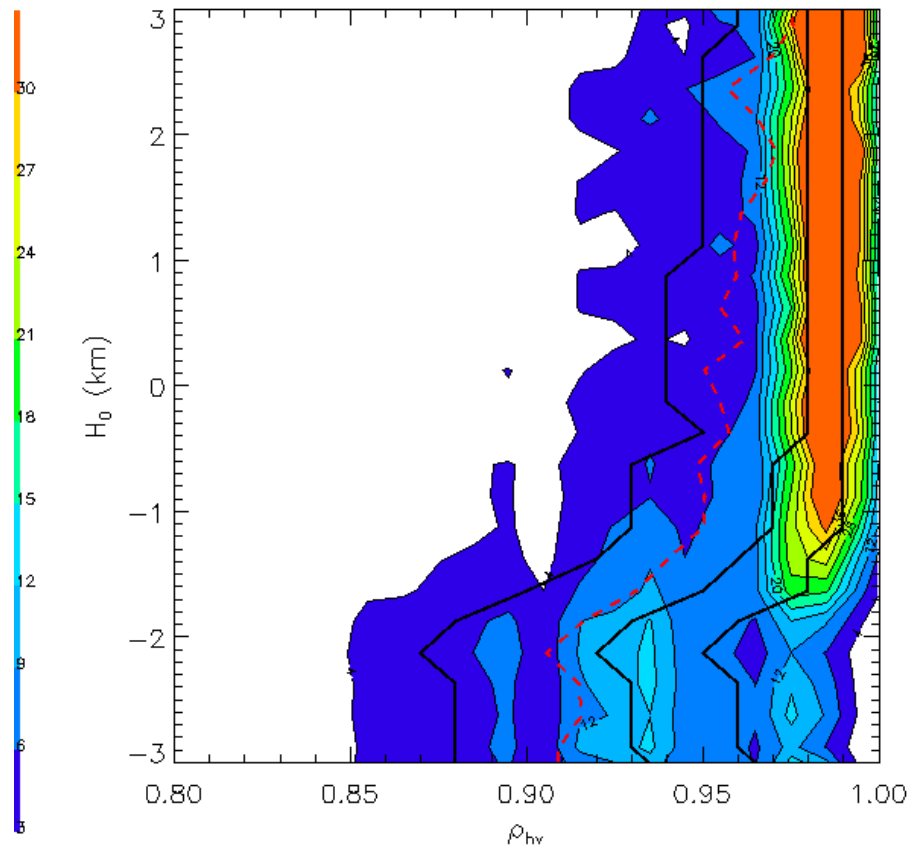
**Figure 5.** Contoured frequency by altitude diagram (CFAD) of  $Z_{dr}$  for the April 10, 2009 ARMOR data set. All data are located within 1 km of a hail report from the April 10, 2009 data set for values of  $Z_h > 45$  dBZ. The  $Z_{dr}$  data are in 0.5 dB bins and the height data are in 250 m bins. The thick black lines from left to right are the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles. The red dashed line is the mean. The height of the 0° C level is indicated (3.0 km).



**Figure 6.** Composite joint relative frequency histogram in percent of ARMOR  $Z_{dr}$  versus  $Z_h$  at low levels for all hail events in [Table 1](#). All data are taken from the lowest 1 km in height above the surface and from within 1 km horizontal distance of a hail report from the composite data set of all nine events for values of  $Z_h > 45$  dBZ. The  $Z_h$  data are in 2 dB bins and the  $Z_{dr}$  data are in 0.5 dB bins. The thick black lines from bottom to top are the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles. The dashed red line is the mean.



**Figure 7.** Contoured frequency by altitude diagram (CFAD) expressed in percent of ARMOR  $Z_{dr}$  (dB) versus height relative to the 0 °C level ( $H_0$ , km) for all hail events in [Table 1](#). All data are located within 1 km of a hail report from the composite data set from nine events for values of  $Z_h > 45$  dBZ. The  $Z_{dr}$  data are in 2 dB bins and the height data are in 250 m bins. The thick black lines from left to right are the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles. The red dashed line is the mean.



**Figure 8.** Contoured frequency by altitude diagram (CFAD) expressed in percent of ARMOR  $\rho_{hv}$  versus height relative to the 0 °C level ( $H_0$ , km) for all hail events in [Table 1](#). All data are located within 1 km of a hail report from the composite data set from nine events for values of  $Z_h > 45$  dBZ. The  $\rho_{hv}$  data are in 0.01 bins and the height data are in 250 m bins. The thick black lines from left to right are the 25<sup>th</sup>, 50<sup>th</sup>, and 75<sup>th</sup> percentiles. The red dashed line is the mean.